ALPO COMET NEWS



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C/2021 A1 (Leonard) closing in on globular cluster M3 in Canes Venatici. Gregg Ruppel took this image on the morning of 2021 December 2 with an ASA 10N f/3.7 and STL11000M from Dark Sky, New Mexico. Image is a composite LRGB with 60 minutes of total exposure.

Table of Contents

SUMMARY	3
APERTURE CORRECTIONS TO MAGNITUDE MEASUREMENTS	3
COMETS CALENDAR FOR NOVEMBER 2021	4
COMETS BRIGHTER THAN MAGNITUDE 10	5
C/2021 A1 (LEONARD)	5
8P/Tuttle	12
67P/Churyumov-Gerasimenko	13
C/2019 L3 (ATLAS)	15
19P/Borrelly	17
COMETS BETWEEN MAGNITUDE 10 AND 13	
4P/Faye	19
6P/d'Arrest	21
29P/Schwassmann-Wachmann	22
57P/du Toit-Neujmin-Delporte	23
C/2017 K2 (PANSTARRS)	24
NEW DISCOVERIES, RECOVERIES AND OTHER COMETS NEWS	26
RECENT MAGNITUDES CONTRIBUTED TO THE ALPO COMETS SECTION	

The monthly ALPO Comet News PDF can be found on the ALPO Comets Section website (<u>http://www.alpo-astronomy.org/cometblog/</u>). A shorter version of this report is posted on a dedicated Cloudy Nights forum (<u>https://www.cloudynights.com/topic/801076-alpo-comet-news-for-december-2021/</u>). All are encouraged to join the discussion over at Cloudy Nights. The ALPO Comet Section welcomes all comet related articles, observations, images, drawings, magnitude estimates, or spectra. One does not have to be a member of ALPO to submit material, though membership is encouraged.

Please send your observations to the Comets Section at $< \underline{comets@alpo-astronomy.org} >$, Coordinator Carl Hergenrother $< \underline{carl.hergenrother@alpo-astronomy.org} >$ and/or Acting Assistant Coordinator Michel Deconinck $< \underline{michel.deconinck@alpo-astronomy.org} >$.

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Summary

It is now or never for C/2021 A1 (Leonard). Currently magnitude 6.5, Leonard has the potential to become a few magnitudes brighter as it passes 0.23 au from Earth on December 12. Complicating any forecast of how bright or observable it might get are a very low solar elongation (down to 15 degrees), uncertainty in the effect of dust forward scattering to enhance Leonard's brightness, and signs that the comet is fading or perhaps even breaking up. This is an object well worth watching in the morning sky from the northern hemisphere during the first half of the month and in the evening sky during the second half of December (though at that time it will be much easier to observe from the southern hemisphere).

Other comets brighter than 10th magnitude in December include C/2019 L3 (ATLAS), 19P/Borrelly, and 67P/Churyumov-Gerasimenko.

In addition to observations submitted directly to the ALPO, we occasionally use data from other sources to augment our analysis. We would like to acknowledge with thanks observations submitted directly to the ALPO as well as those originally submitted to the International Comet Quarterly, Minor Planet Center, and COBS Comet Observation Database. We would also like to thank the Jet Propulsion Laboratory for making available their Small-Body Browser and Orbit Visualizer and Seiichi Yoshida for his Comets for Windows programs that is used to produce the lightcurves in these pages. And last but not least, we'd like to thank <u>Syuichi Nakano</u> and the Minor Planet Center for their comet orbital elements, the asteroid surveys and dedicated comet hunters for their discoveries, and all of the observers who volunteer their time to adding to our knowledge of these amazing objects.

Aperture Corrections to Magnitude Measurements

We try to include up to date lightcurves for most of the objects discussed in this report as well as applying aperture corrections to the visual observations. All magnitude estimates are affected by many factors including instrumental (aperture, focal length, magnification, type of optics), environmental (sky brightness due to moonlight, light pollution, twilight, aurora activity, zodiacal light, etc), cometary (degree of condensation, coma color, strength and type of gas emission lines, coma-tail interface) and personal (sensitivity to different wavelengths, personal technique, observational biases). The correction used here only corrects for differences in aperture [C. S. Morris, On Aperture Corrections for Comet Magnitude Estimates. Publ Astron Soc Pac 85, 470, 1973]. Visual observations are corrected to a standard aperture of 6.78 cm by 0.019 magnitudes per centimeter for refractors and 0.066 magnitudes per centimeter for reflectors. As our work develops, we will investigate the determination of personal corrections for each observer for each individual comet as well as for digital observations.

Comets Calendar for December 2021

Dec 02-03	– C/2021 A1 (Leonard) passes ~0.1 deg of 6 th mag globular cluster M3
Dec 02	-430P/Scotti at perihelion (q = 1.55 au, 5.5-year orbit, V ~ 17, 2 nd observed return, discovered
	in 2011, missed at 2016 return)
Dec 04	– New Moon
Dec 06	- 19P/Borrelly orbit plane crossing
Dec 08	– C/2021 A1 (Leonard) orbit plane crossing
Dec 08	 - 436P/Garradd at perihelion (q = 1.96 au, 14.4-year orbit, V ~ 19, 2nd observed return, discovered in 2007)
Dec 10	– First Quarter Moon
Dec 12	– C/2021 A1 (Leonard) passes ~0.2 deg of 9 th mag globular cluster NGC 6366
Dec 12	 C/2021 A1 (Leonard) passes 0.2334 au (34.9 million km, 21.7 million miles) from Earth (Dec 12 – 13:53 UT)
Dec 14	- C/2021 A1 (Leonard) passes across bright emission nebula M16, the Eagle Nebula
Dec 14	- 402P/LINEAR at perihelion (q = 3.94 au, 18.6-year orbit, V ~ 16-17, 2^{nd} observed return, discovered in 2003)
Dec 16	- 173P/Mueller at perihelion (q = 4.22 au, 13.6-year orbit, V ~ 19, very asymmetric lightcurve with peak activity 1-2 years BEFORE perihelion, peaked at V~16, 3^{rd} observed return)
Dec 16	-P/2021 R1 (PANSTARRS) at perihelion (q = 4.89 au, 24.4-year orbit, V ~ 20, discovered last
	September)
Dec 18	- C/2021 A1 (Leonard) passes 0.0285 au (4.26 million km, 2.65 million miles) from Venus (Dec 18 - 02:06 UT), will appear 5.1 deg from Venus as seen from Earth
Dec 18	– Full Moon
Dec 18	- 221P/LINEAR at perihelion (q = 1.75 au, 6.4-year orbit, V ~ 19, currently in solar conjunction, 4^{th} observed return, discovered in 2002)
Dec 19	– C/2017 K2 (PANSTARRS) orbit plane crossing
Dec 21	-C/2021 U4 (Leonard) at perihelion (q = 1.79 au, ~300-year orbit, V ~ 20-21)
Dec 21	-C/2021 R2 (PANSTARRS) at perihelion (q = 7.31 au, ~110,000-year orbit, V ~ 20)
Dec 21	– 8P/Tuttle orbit plane crossing
Dec 26	– Last Quarter Moon
Dec 30	- 395P/Catalina-NEAT at perihelion (q = 4.06 au, 16.8-year orbit, V ~ 17-18, 2 nd observed return, discovered in 2005)
Dec 30 Ian 1	$4P/E_{2}$ passes between 9^{th} mag open cluster NGC 2254 and emission network IC 448

Dec 30-Jan 1 – 4P/Faye passes between 9th mag open cluster NGC 2254 and emission nebula IC 448

C/2021 A1 (Leonard)

Discovered 2021 January 3 by Greg Leonard of the Catalina Sky Survey with the 1.5-m on Mount Lemmon

Orbit (from Syuichi Nakano, personal email)

C/2021 A1 (Leonard) Epoch 2022 Jan. 21.0 TT = JDT 2459600.5 T 2022 Jan. 3.29809 TT Nakano q 0.6152601 (2000.0) P Q z -0.0000451 Peri. 225.09163 +0.63774181 +0.29161284 +/-0.0000011 Node 255.89505 +0.72791572 -0.53079785 e 1.0000277 Incl. 132.68632 -0.25184139 -0.79574845 From 1274 observations 2020 Apr. 11-2021 Nov. 18, mean residual 0".57. (1/a) org.= +0.000501, (1/a) fut.= -0.000105 (+/-0.000001), Q= 8.

Comet Magnitude Formula (from ALPO and COBS data)

 $m1 = 11.4 + 5 \log d + 5.7 \log r$ [T-325 to T-240 days, where T = date of perihelion] m1 = 7.3 + 5 log d + 12.5 log r [T-240 to T-47 days] m1 = 8.3 + 5 log d - 3.8 log r [from T-47 days and onward?]

Recent Magnitude Measurements Contributed to the ALPO Comets Section

Recent Magnitude Measurements in ICQ format: Comet Des YYYY MM DD.DD Mag SC APER FL POW COMA TAIL ICO CODE Observer Name (UT) T Dia DC LENG PA 2021A1 2021 12 02.49 Z 6.5 U4 7.2R 5a600 26.3 > 1.0 313 ICQ xx HER02 Carl Hergenrother 2021A1 2021 12 02.48 M 6.7 TK 5.0B 10 8 5/ 0.7 320 ICQ xx HER02 Carl Hergenrother

 2021A1
 2021 12
 01.46
 M
 0.7
 TK
 5.08
 10
 6
 5/
 0.3
 320
 ICQ xX
 HER02
 Call Hergenrother

 2021A1
 2021 12
 01.47
 M
 7.0
 TK
 5.08
 10
 6
 5/
 0.3
 320
 ICQ xX
 HER02
 Call Hergenrother

 2021A1
 2021 12
 01.17
 I
 6.8
 TK
 12.68
 5
 25
 10
 6
 45.0m300
 ICQ XX
 HER02
 Carl Hergenrother

 2021A1
 2021 11
 30.47
 M
 7.1
 TK
 5.08
 10
 5
 0.3
 320
 ICQ xX
 HER02
 Carl Hergenrother

 2021A1
 2021 11
 28.50
 M
 7.2
 TK
 5.08
 10
 7
 5
 0.2
 315
 ICQ xX
 HER02
 Carl Hergenrother

 2021A1
 2021 11
 28.18
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 25.0m300
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 Michel Deconinck

 2021A1
 2021 11
 26.46
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 7.9< 2021A1 2021 11 23.15 S 8.4 TK 32.0L 5 80 4 5/ 5.0 310 ICQ XX PILO1 Uwe Pilz

 2021A1
 2021 11
 18.24
 S
 7.9
 TK
 10.08
 25
 8
 3
 0.15
 320
 ICQ
 XX
 GON05
 Juan
 Jose
 Gonzalez
 Suarez

 2021A1
 2021
 11
 17.51
 S
 8.2
 TK
 5.0B
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 7
 4
 ICQ
 XX
 GON05
 Juan
 Jose
 Gonzalez
 Suarez

 2021A1
 2021
 11
 17.50
 V
 8.5
 U4
 10.6R
 5a600
 13.4
 20
 m320
 ICQ
 xx
 HER02
 Carl
 Hergenrother

 2021A1
 2021
 11
 16.20
 I
 8.4
 TK
 25.0C10
 62
 2.5
 7
 5.0m280
 ICQ
 XX
 DECaa
 Michel
 Deconinck

 2021A12021 1115.48V9.0U410.6R5a60012.82021A12021 1115.52S8.2TK5.0B1054 12 m320 ICQ xx HER02 Carl Hergenrother

 2021A1
 2021 11
 05.23
 S
 9.7 TK 20.3T10
 77
 5
 3/
 0.3
 330
 ICQ XX GON05
 Juan Jose Gonzalez Suarez

 2021A1
 2021 11
 04.49
 S
 10.0 TK 12.5B
 30
 4
 ICQ XX HER02
 Carl Hergenrother

 2021A1
 2021 11
 02.15
 S
 10.3 TI 29.8L 4
 79
 3
 4
 6.0m340
 ICQ XX HAR11
 Christian Harder

As of December 2-3, C/2021 A1 (Leonard) is around magnitude 6.5, bright enough that some observers have reported it visible to the naked eye. Visual observers have also been observing a tail up to 0.75 degrees in length. Imagers have noted some interesting coma features, which unfortunately are sometimes associated with splitting or disruption events.

Eleven months ago, C/2021 A1 (Leonard) was discovered at 19th magnitude by Catalina Sky Survey astronomer Greg Leonard with the Mount Lemmon 1.5-m reflector when the comet was a distant 5.1 au from the Sun. The comet could have easily been discovered months earlier as pre-discovery observations from Mount Lemmon and PANSTARRS were found back to April 2020 when the comet was 7.5 au from the Sun. With a relatively small perihelion distance of 0.62 au in 2022 January and close approach to Earth on December 12 at 0.233 au (34.9 million km, 21.7 million miles), there was some excitement that Leonard could become a bright object.

The story of Comet Leonard has seen a few twists and turns. Based on the Pan-STARRS and Mount Lemmon photometry submitted to the Minor Planet Center, Leonard rapidly brightened throughout 2020 at a $2.5n \sim 11.7$ rate. Between January and June 2021, that rapid brightening slowed significantly to a sluggish $2.5n \sim 5.6$ rate suggesting activity was barely increasing beyond steady state. Six months ago, Leonard's prospects weren't looking good, but in July a new rapid brightening phase commenced. A fit to the photometry shows two possible brightening "legs" ($2.5n \sim 20.6$ between July and early October and $2.5n \sim 12.5$ between early October and mid-November). Regardless of the exact fits or the number of "legs", the comet brightened from apparent magnitude 17 to 8 and excitement started to build again.

Then something changed in mid-November. Since about November 16, Leonard has not only ceased its rapid brightening, but has faded in an absolute sense with a fading trend of $2.5n \sim -4$. Leonard has faded by ~ 2 magnitudes after accounting for changes in heliocentric and geocentric distance. That is a dramatic decrease in activity for an inbound comet and suggests C/2021 A1 may not be healthy.



Figure 1 - Brightness evolution of C/2021 A1 (Leonard) since the start of 2020. The comet photometry has been corrected foraperture, phase angle effects as well as normalized to 1 au from the Earth and Sun. Data includes CCD/CMOS and visual photometry submitted to the ALPO and data submitted to the COBS site by Michael Lehmann.



Figure 2 – Same data as above except plotted as aperture corrected apparent magnitude versus date. Plot produced in Seiichi Yoshida's Comets for Windows.



Figure 3 - Same as except only showing the most recent aperture corrected apparent magnitudes.

You may be wondering how a comet could intrinsically fade by two magnitudes while its apparent magnitude brightened from magnitude 8.2 to 6.8 over the same period. The reason is that between November 16 and December 3 Leonard's distance to the Sun has fallen from 1.16 to 0.90 au and its distance to Earth from 1.06 to 0.45 au. Its phase angle has also increased from 52 to 86 degrees though that change should have only produced ~ 0.1 magnitudes of fading.

If you are having a sense of deja vue, you're not alone. We saw a similar series of events in 2020 with C/2019 Y4 (ATLAS). Comet ATLAS was predicted to be a bright object but after brightening to ~7th magnitude about 2.5 months before perihelion, it's brightening stalled and then enter a slow fade. A few weeks after the brightening stall, the nucleus was observed to have split into several smaller components. Could we be seeing the same with Leonard? Possibly. Leonard brightness stall and fading is similar to that of C/ATLAS. So far, we haven't seen any sign of multiple nuclei, but then again it took a few weeks for ATLAS' secondary nuclei to become visible.

So, for the next few paragraphs, let's (mostly) focus on the things we can predict. Leonard starts the month at 0.93 au from the Sun and 0.52 au from Earth. On December 12, a minimum Earth-comet distance of 0.23 au is reached with the comet located 0.77 au from the Sun. Not only is the Earth-comet distance rapidly changing, but due to the comet passing Earth on its sunward side, Leonard's phase angle increases from 80 degrees on the 1st to 160.5 degrees on the 14th. Such a large phase angle could produce up to 3 magnitudes of extra brightness due to forward scattering by small dust in the comet's coma and tail. The exact amount of forward scattering remains to be seen and depends on the comet's dust-to-gas ratio and whether it is still producing fine sized dust at that time, something that could be in question if it has disintegrated.

Leonard is a morning object only visible from the northern hemisphere as the month begins though it is rapidly diving towards the Sun. On December 2-3, it will pass through the outer region of the bright 6th magnitude globular cluster M3. By December 12, the comet will too close to the Sun to be observed outside of astronomical twilight. The next night, the comet shifts into the evening sky and will be within 18 degrees of the Sun but only a degree above the horizon by the end of nautical twilight. Unless the comet is bright (1st-2nd magnitude) and condensed (like last year's NEOWISE) it may not be visible against such a bright sky. Unfortunately, the bright sky may prevent observation of Leonard photobombing two more deep sky objects: on December 12 is passes ~0.2 deg from 9th mag globular cluster NGC 6366 and on December 14 it passes over the bright emission nebula M16, the Eagle Nebula. How often do you see a comet pass over two Messier objects in an apparition, let alone in 12 days? The southern hemisphere finally gets their chance to observe the comet starting on December 17-18. The 18th also witnesses an extremely close approach to Venus of 0.0285 au (4.26 million km, 2.65 million miles). Here on the Earth the two will appear about 5.1 degrees apart.

The southern hemisphere will have the best views during the remainder of the month. While still visible from the northern hemisphere it will be a horizon hugger and only a few degrees above the horizon before the start of astronomical twilight. By the end of December, its distance to the Earth will have increased back to 0.83 au while its distance to the Sun continues to drop as it approaches perihelion on January 3 at 0.62 au. This month sees the comet move through the constellations of Canes Venatici (Dec 1-3), Boötes (3-8), Serpens (8-10), Hercules, (10-11), Ophiuchus (11-14), Scutum (14-15), Sagittarius (15-20), Microscopium (20-29), and Pisces Austrinus (29-31).

So now the difficult, or even foolish, discussion... how bright will Leonard get? We have a few complications: will the comet continue to intrinsically fade, or even disintegrate, and how much of an effect will dust forward scattering have? There are several scenarios that can play out. One could expect a resumption of the brightening trend and a 3-magnitude enhancement due to forward scattering. That would result in a 1st magnitude comet and perhaps visible even when at small elongations. Then again, if the fading continues and there is little to no

forward scattering, a peak around 5-6th magnitude would result in Leonard being much too faint to be seen when located close to the Sun. The ephemeris below presents two scenarios: the magnitude in the 'Mag NoFS' column assumes the fading continues and there is no enhancement due to forward scattering (peak brightness of 5.6), the 'Mag FS' column also assumes the fading continues but with a maximum of ~3 magnitudes of forward scattering enhancement for a peak brightness of 2.0. I guess we'll have to wait and see what happens.

C/2021 A1 (Leonard)

C/202	21 A1	L (L	eona	ard)														Ma	ax El	L
																		((deg)	
I	Date		R.	Α.	Dec	21.	1	r		d	PhAno	g	Elong	Мас	g	Mag	4	ЛС		10S
														Nol	FS	FS	Ast	Nau	Ast	Nau
2021	Dec	01	13	21	+29	58	0.9	931	0	.522	80		68M	7.(С	7.0	49	55	0	0
2021	Dec	02	13	30	+29	17	0.9	917	0	.487	83		67M	6.9	9	6.9	47	54	0	0
2021	Dec	03	13	40	+28	27	0.9	902	0	.453	86		66M	6.8	8	6.8	46	52	0	0
2021	Dec	04	13	52	+27	24	0.8	888	0	.419	90		64M	6.0	6	6.6	44	50	0	0
2021	Dec	05	14	05	+26	05	0.8	874	0	.387	94		62M	6.5	5	6.4	41	47	0	0
2021	Dec	06	14	21	+24	26	0.8	859	0	.356	99		59M	6.3	3	6.2	38	44	0	0
2021	Dec	07	14	39	+22	20	0.8	846	0	.326	105		55M	6.1	1	5.8	34	40	0	0
2021	Dec	80	14	59	+19	39	0.8	332	0	.299	112		51M	6.0	С	5.6	29	35	0	0
2021	Dec	09	15	23	+16	17	0.8	818	0	.276	119		46M	5.8	8	5.2	22	29	0	0
2021	Dec	10	15	49	+12	06	0.8	805	0	.256	128		39M	5.	7	4.7	15	21	0	0
2021	Dec	11	16	19	+07	80	0.	792	0	.242	137		32M	5.0	6	4.1	7	13	0	0
2021	Dec	12	16	50	+01	32	0.	779	0	.235	147		25M	5.0	6	3.4	0	4	0	0
2021	Dec	13	17	22	-04	19	0.	767	0	.234	155		18E	5.0	6	2.6	0	1	0	0
2021	Dec	14	17	55	-09	58	0.	755	0	.241	160		15E	5.	7	2.0	0	3	0	0
2021	Dec	15	18	25	-15	01	0.	743	0	.254	159		15E	5.8	8	2.3	0	4	0	0
2021	Dec	16	18	53	-19	15	0.	731	0	.273	154		18E	6.0	С	3.1	0	5	0	0
2021	Dec	17	19	18	-22	40	0.	720	0	.297	147		22E	6.2	2	4.0	1	6	0	4
2021	Dec	18	19	39	-25	22	0.	709	0	.324	141		26E	6.4	4	4.7	2	6	1	9
2021	Dec	19	19	58	-27	28	0.0	699	0	.354	135		29E	6.0	6	5.3	2	7	4	12
2021	Dec	20	20	14	-29	06	0.0	689	0	.387	130		32E	6.9	9	5.8	3	7	7	15
2021	Dec	21	20	28	-30	22	0.0	680	0	.421	125		34E	7.1	1	6.2	3	7	10	18
2021	Dec	22	20	40	-31	23	0.0	671	0	.456	120		36E	7.3	3	6.6	3	7	11	19
2021	Dec	23	20	50	-32	12	0.0	663	0	.492	116		37E	7.4	4	6.9	3	7	13	21
2021	Dec	24	20	58	-32	51	0.0	655	0	.529	111		38E	7.0	6	7.2	3	7	14	22
2021	Dec	25	21	05	-33	24	0.0	648	0	.566	107		38E	7.8	8	7.5	3	7	14	23
2021	Dec	26	21	12	-33	51	0.0	641	0	.604	104		39E	7.9	9	7.7	3	7	15	23
2021	Dec	27	21	17	-34	13	0.0	636	0	.642	100		39E	8.2	1	7.9	3	7	15	24
2021	Dec	28	21	21	-34	32	0.0	631	0	.680	97		39E	8.2	2	8.1	2	6	16	24
2021	Dec	29	21	25	-34	49	0.0	626	0	.718	93		39E	8.4	4	8.4	2	6	16	24
2021	Dec	30	21	29	-35	03	0.0	622	0	.757	90		39E	8.5	5	8.5	2	6	16	24
2021	Dec	31	21	32	-35	15	0.0	620	0	.795	87		38E	8.0	6	8.6	1	5	16	24

Let's again assume a healthy normal comet for the remainder of the month, what will the tails of Leonard look like? Currently, the comet's dust and gas tails are superimposed on each other. Using the Finson-Probstein analysis program of the Comet Toolbox (https://www.comet-toolbox.com/FP.html), we should expect that to continue for another week and a half. With the Earth crossing the comet's orbital plane on December 8, the dust tail may become very narrow for a few days. After close approach, the dust tail should diverge from the antisolar direction and fan out away from the gas tail.

One must be careful when interpreting the following figure. It only shows the possible orientation of the tails relative to the nucleus. It should not be used to infer the absolute or relative lengths of the tails. All bets are off if the comet continues to fade or even completely disrupt.

Here to a bunch of lost sleep this month watching C/Leonard!



Figure 4 - Finson-Probstein analysis for C/2021 A1 (Leonard) as modeled with the Comet Toolbox (<u>https://www.comet-toolbox.com/FP.html</u>). Synchrones are lines of constant time of dust release (in days prior to the modeled data) while syndynes are lines of constant dust size. The values of the syndynes are the "beta" of the particles which is the ratio of solar radiation pressure to solar gravity and is inversely proportional to grain size.



Figure 5 - Sketches of C/2021 A1 (Leonard) by [top] Ludovic Perbet on November 30 and [bottom] Michel Deconinck on December 1.



Figure 6 - Recent image of C/2021 A1 (Leonard) by Martin Mobberley.

Discovered on 1790 January 9 by Pierre F. A. Mechain Rediscovered on 1858 January 5 by Horace Tuttle Solution Minor Planet Center, MPEC 2021-W138) ************************************	1											
Obti (from Minor Planet Center, MPEC 2021-W138) 8P/Tuttle Epoch 2022 Jan. 21.0 TT = JDT 2459600.5 7 2021 Aug. 27.73567 TT Rudenko 9 1.0259957 2000.0) P 0.07229128 Peri. 207.48835 -0.26845021 -0.13638410 e 0.8202233 Incl. 54.91130 +0.00597787 -0.85030055 P 13.6 From 271 observations 2008 Feb. 12-2021 Nov. 23, mean residual 0".6. Nongravitational parameters A1 = +0.01, A2 = -0.0091. Ephemerides (produced with Selichi Yoshida's Comets for Windows program) 8P/Tuttle Max E1 (deg) Date R.A. Decl. r d Color 0 14 19 - 49 21 1.707 2.381 2021 Dec 01 14 19 - 49 21 1.707 2.381 37M Lup 11.8 19 2021 Dec 11 14 53 - 51 04 1.810 2.470 38M Lup 12.2 0 20 2021 Dec 11 14 53 - 51 45 1.862 2.512 39M Lup 12.8 21 2021 Dec 21 15	Discovered on 1790 January 9 by Pierre F. A. Mechain Rediscovered on 1858 January 5 by Horace Tuttle											
BP/Tuttle Epoch 2022 Jan. 21.0 TT = JDT 2459600.5 T 2021 Jau, 27.73567 TT Rudenko q 1.0259957 (2000.0) P Q n 0.07229128 Peri. 207.48835 -0.26845021 -0.50831915 a 5.7070558 Node 270.20165 +0.96327501 -0.13638410 e 0.8202233 Incl. 54.91130 +0.00597787 -0.85030055 P 13.6 From 271 observations 2008 Feb. 12-2021 Nov. 23, mean residual 0".6. Nongravitational parameters Al = +0.01, A2 = -0.0091. Ephemerides (produced with Seüchi Yoshida's Comets for Windows program) Max El (deg) Date R.A. Decl. r d Elong Const Mag 40N 40S 2021 Dec 06 14 19 -49 21 1.707 2.381 2021 Dec 11 14 9 -9 21 1.707 2.487 2021 Dec 11 14 53 -51 04 1.862 2.512 39M Lup 12.5 0 2021 Dec 11 14 53 -51 45 1.862 2.512 39M Lup 12.8 0 21 2021 Dec 16 15 10 -51 45 1.862 2.512 39M Lup 13.1 0 22 2021 Dec 26 1	Orbit (from Minor Planet Center, MPEC 2021-W138)											
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<pre>m1 = 7.0 + 5 log d + 20 log r(t-25) [Ref: Seiichi Yoshida] Magnitude Measurements Submitted to the ALPO Comets Section Recent Magnitude Measurements in ICQ format: Comet Des YYYY MM DD.DD Mag SC APER FL POW COMA TAIL ICQ CODE Observer Name</pre>	Comet Magnitude Formula											
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Recent Magnitude Measurements in ICQ format: Comet Des YYYY MM DD.DD Mag SC APER FL POW COMA TAIL ICQ CODE Observer Name (UT) T Dia DC LENG PA 8 2021 11 13.71 xS 10.4 TK 40.0L 4 59 3.5 3 ICQ XX WYA Christopher Wyatt	Magnitude Measurements Submitted to the ALPO Comets Section											
	Recent Magnitude Measurements in ICQ format: Comet Des YYYY MM DD.DD Mag SC APER FL POW COMA TAIL ICQ CODE Observer Name (UT) T Dia DC LENG PA 8 2021 11 13.71 xS 10.4 TK 40.0L 4 59 3.5 3 ICQ XX WYA Christopher Wyatt											

A large geocentric distance, small solar elongation, and invisibility from the northern hemisphere have limited observations of 8P/Tuttle during the current return. Perihelion was back on 2021 August 27 with a peak brightness of 8.5-9.0 during September. Last month Chris Wyatt visually observed Tuttle on November 13 with a 0.4-m reflector at 59x. He measured a brightness of 10.4 with a weakly condensed (DC = 3) 3.5' coma.

The comet will be rapidly fading in December from around magnitude ~ 12 to ~ 14 as it moves through Lupus (Dec 1-22) and Norma (22-31). As has been the case for much of this apparition, Tuttle will only be visible to southern hemisphere observers.

Imagers are encouraged to take deep, wide field images of Tuttle during the 2nd half of the December as the Earth will be crossing the Tuttle's orbital plane on December 21. The dust trail consists of larger dust particles orbiting along the comet's orbit. These particles are populating the Ursid meteor stream, so any image of Tuttle's dust trail may be an image of future Ursids.

This will probably be the last month we report on 8P as it will fainter than our 13th magnitude cutoff next month. Looking towards the future, two returns from now will be much better with Tuttle passing 0.18 au from Earth on 2048 December 28 and brightening to 4th magnitude.

67P/Churyumov-Gerasimenko



67P was discovered on photographic plates taken on 1969 September 11 by Kiev University Astronomical Observatory astronomers Klim Ivanovic Churyumov and Svetlana Ivanovna Gerasimenko working at the Alma-Ata Astrophysical Institute in current day Kazakhstan. The current apparition is 67P's 9th observed return with perihelion back on 2021 November 2 at 1.21 au. A close approach to Earth at 0.42 au on November 12 makes this the comet's best return since 1982 when it came marginally closer to Earth at 0.39 au. Though this return

will be the best for the remainder of the century, it will make similar, but slightly larger, approaches to Earth in 2034 (0.45 au), 2067 (0.44 au), and 2080 (0.49 au).

Last month, visual observers found 67P to be between magnitude 9.0 and 9.9 (aperture corrected to 8.6 to 9.6). with up to a quarter degree of tail. The comet spends December in the morning sky in Cancer. Based on previous apparitions, 67P should be at its brightest during the first half of December at around magnitude 8.3 and slightly fade to around 8.6 by New Year's.



Figure 7 – Image of 67P taken by Chris Schur on November 13 with a 10" f/3.9 Orion Astrograph Newtonian.



Comet 67P (Churyumov-Gerasimenko) Takahashi Mewlon 10" f10 - 195x

2021/11/07 - 05h05 UTC F.O.S.: 20'

ICQ: 67 2021 11 07.21 B 9.6 TK 25.0C10 195 1.8 7 2.2m260

Aquarellia Observatory

Figure 8 - Sketch of 67P by Michel Deconinck on Nov 7 with a Takahashi Mewlon 0.25-m at 195x.

C/2019 L3 (ATLAS)

Discovered 2019 June 10 by the ATLAS survey with one of their 0.5-m f/2 Schmidt Dynamically old long-period comet

Orbit (from Syuichi Nakano, personal email)

C/2019 L3 (ATL	AS)			
Epoch 2022 Jan. 21	.0 TT =	JDT 2459600	.5	
T 2022 Jan. 9.619	30 TT			Rudenko
q 3.5545066		(2000.0)	P	Q
z -0.0004534	Peri.	171.61066	-0.26052094	-0.66630823
+/-0.000003	Node	290.79019	+0.83675993	+0.20517882
e 1.0016115	Incl.	48.36122	+0.48162398	-0.71689259
From 2893 observat	ions 20	19 June 10-2	021 Nov. 19, mean	residual 0".4
1/a(orig) = +0.000	102 AU*	*-1, 1/a(fut	$) = -0.000881 \text{ AU}^{*}$	*-1.

Ephemerides (produced with Seiichi Yoshida's Comets for Windows program)

C/2019 L3 (A	ATLAS)							Max (d	El eq)
Date	R.A.	Decl.	r	d	Elong	Const	Mag	40N	40s
2021 Dec 01	07 40	+37 31	3.573	2.798	135M	Lyn	9.1	87	12
2021 Dec 06	07 37	+36 55	3.568	2.746	141M	Lyn	9.0	87	13
2021 Dec 11	07 33	+36 15	3.564	2.700	146M	Lyn	9.0	86	14
2021 Dec 16	07 29	+35 33	3.561	2.661	152M	Aur	8.9	85	15
2021 Dec 21	07 24	+34 48	3.559	2.629	157M	Gem	8.9	85	15
2021 Dec 26	07 19	+33 59	3.557	2.605	162M	Gem	8.9	84	16
2021 Dec 31	07 14	+33 07	3.555	2.589	167M	Gem	8.9	83	17
2022 Jan 05	07 08	+32 12	3.554	2.581	170M	Gem	8.9	82	18

Comet Magnitude Formula and Lightcurve (from ALPO and COBS data)

m1	=	2.0	+	5	log	d	+	12.3	log	r	[through T-550	days; T	. =	date	of	perihelion]
m1	=	-4.6	+	5	log	d	+	20.8	log	r	[T-550 to T-60	days]				
m1	=	2.4	+	5	loq	d	+	8.0	loq	r	[T-60 days and	onwards	5]			



The next two comets also share the near-opposition sky with C/2019 L3 (ATLAS). With only a month to go till its perihelion at 3.57 au (T = 2022 January 9), ATLAS should finally reach peak brightness around magnitude 9. In November, Chris Wyatt, Carl Hergenrother, Christian Harder, and J. J. Gonzalez made 6 visual observations finding the comet between magnitude 9.5 and 10.2 (aperture corrected range of 9.2 and 9.8). Chris and Christian were also able to visually detect a tail up to 6' in length in 0.3- to 0.4-m telescopes.

C/2019 L3 (ATLAS) will be approaching opposition while moving through Lynx (Dec 1-14), Auriga (14-17), and Gemini (17-31) in the morning sky. While well placed for northern observers, it is also visible but low from the southern hemisphere.



Figure 9 - C/2019 L3 (ATLAS) cruised past galaxy PGC 21754 (just above the comet's head) and interacting galaxies NGC 2444 and 2445 (near upper left corner of image). Eliot Herman took this image with the iTelescopes T19 0.43-m reflector on November 16. The image consists of a series of 300 sec exposures co-added to total 30 min in Luminance and 15 min in H-alpha filters.

Discovered 1904 December 28 by the Alphonse Borrelly Short-period comet with orbital period of ~6.85 years

Orbit (from Minor Planet Center, MPEC 2021-W138)

	19P/Borrelly				
Εŗ	ooch 2022 Jan. 2	21.0 TT =	JDT 2459600.	5	
Т	2022 Feb. 1.830	20 TT			Rudenko
q	1.3062541		(2000.0)	P	Q
n	0.14401007	Peri.	351.92096	+0.38674585	-0.79279248
а	3.6047416	Node	74.24710	+0.87109441	+0.14639270
е	0.6376289	Incl.	29.30470	+0.30269155	+0.59164961
Ρ	6.84				

From 480 observations 2015 Jan. 11-2021 Nov. 29, mean residual 0".7. Nongravitational parameters A1 = -0.86, A2 = -0.7759.

Ephemerides (produced with Seiichi Yoshida's Comets for Windows program)

19	P/Borre	lly	7									Max	El
												(d	leg)
	Date		R.A	Α.	Decl	•	r	d	Elong	Const	Mag	40N	40S
20	21 Dec	01	23 2	26	-38 C)7	1.500	1.175	87E	Gru	10.3	12	63
20	21 Dec	06	23 3	33	-35 1	. 4	1.473	1.174	85E	Scl	10.1	15	58
20	21 Dec	11	23 4	11	-32 1	.2	1.447	1.173	83E	Scl	9.9	18	54
20	21 Dec	16	23 4	19	-29 C)3	1.423	1.174	82E	Scl	9.7	21	49
20	21 Dec	21	23 5	58	-25 4	17	1.401	1.175	80E	Scl	9.5	24	44
20	21 Dec	26	00 0)7	-22 2	24	1.381	1.178	78E	Cet	9.4	27	40
20	21 Dec	31	00 1	16	-18 5	66	1.363	1.183	77E	Cet	9.2	31	36
20	22 Jan	05	00 2	26	-15 2	23	1.347	1.189	76E	Cet	9.1	33	32

Comet Magnitude Formula (from Seiichi Yoshida)

 $m1 = 5.5 + 5 \log d + 25.0 \log r$



19P/Borrelly should be one of the better comets of 2022 when it may reach 9th magnitude around the time of its 2022 February 1 perihelion (at 1.31 au). Borrelly is steadily moving north and is now visible in the evening sky from both hemispheres as it moves through Grus (Dec 1), Sculptor (1-22), and Cetus (22-31). While no magnitude estimates were submitted to the ALPO for Borrelly in November, observations submitted to the COBS site place the comet around magnitude 10.5 at the end of November which is line with the prediction above. By the end of the year, the comet should be around magnitude 9.2. Images from the past week show a long dust trail located along the orbit of Borrelly. This feature should become stronger and better defined as the Earth crosses the plane of Borrelly's orbit on December 6.



Figure 10 - Image of 19P/Borrelly taken by Martin Mobberley on December 1 with a iTelescopes Celestron RASA 11". The dust trail along Borrelly's orbit extends from the lower right to upper left and runs through the coma of the comet.



Figure 11 - The Earth crossing the plane of 19P/Borrelly's orbit in early December. Image made with the JPL Small Body Database Orbit Viewer.

4P/Faye



4P/Faye was a visual discovery by Herve Faye (Royal Observatory, Paris, France) on 1843 November 23. The comet was abnormally bright and reported to be visible to the naked eye only days after discovery. It has never rivaled its discovery apparition in brightness and at its best only gets to 9th magnitude (in 1991 and 2006).

This year's apparition is Faye's 22nd observed return with the comet having been missed at its 1903 and 1918 returns. 2021 is a moderately good, but not great, apparition with perihelion on 2021 September 8 at 1.62 au. Even though perihelion was a month ago, the comet will continue to move closer to the Earth until December 5 (0.94 au). As a result, it will stay close to maximum brightness through November. It is a morning object observable from both hemispheres as its moves through Monoceros.

Faye was well observed in October with no less than a dozen visual observations submitted to the ALPO. The most recent observations from November 5th placed the comet around magnitude 10.6-11.2 (aperture corrected to 10.2 to 10.9). While the tail has been a striking feature in images, visual observers have also caught glimpses of the tail. Chris Wyatt reported a 7.5' long tail with a 0.4-m reflector on October 9.

Photo Op: Dec 30-Jan 1 – 4P/Faye passes between 9th mag open cluster NGC 2254 and emission nebula IC 448



Figure 12 - Denis Buczynski caught 4P/Faye on November 9 with a Celestron C14 and ASI1600MM camera. The final image consists of 12x120s coadded image.

Discovered on 1851 June 28 by the Heinrich Ludwig d'Arrest

Orbit (from MPEC 2021-W138)

6P/d'Arres Epoch 2022 Jau T 2021 Sept. 3 q 1.3545380 n 0.1506734 a 3.4976653 e 0.6127308 P 6.54 From 3229 obse Nongravi	t n. 21.0 TT = 17.76118 TT (4 Peri. Node Incl. ervations 198 tational para	JDT 2459600.5 2000.0) 178.08852 138.93495 19.51219 7 Mar. 31-202 meters A1 = -	<pre> P +0.7328913 -0.6285529 -0.2603680 21 Nov. 29, 1 +0.54, A2 = -</pre>	Ruden 3 +0.64 2 +0.76 5 -0.03 mean residu +0.0991.	ko Q 399340 434200 246255 al 1".0.		
Ephemerides (pro	duced with Seiich	<u>i Yoshida's Con</u>	nets for Window	<u>vs program)</u>			
6P/d'Arrest Date 2021 Dec 01 2 2021 Dec 06 2 2021 Dec 11 2 2021 Dec 16 2 2021 Dec 21 2 2021 Dec 26 2 2021 Dec 31 2 2022 Jan 05 0 Comet Magnitude m1 = 6.6 + 5 2	R.A. Decl. 22 30 -27 33 22 45 -26 23 23 00 -25 09 23 14 -23 50 23 28 -22 30 23 41 -21 07 23 54 -19 44 00 06 -18 20 2 Formula (from from from from from from from from	r 1.593 1 1.622 1 1.651 1 1.682 1 1.714 1 1.746 1 1.779 1 1.812 1 t to ALPO and C log r(t-60)	d Elong .423 80E .483 79E .545 78E .609 76E .675 75E .743 73E .813 72E .884 70E COBS data, seas	Const Mag PsA 10.7 PsA 10.9 PsA 11.0 Aqr 11.2 Aqr 11.4 Aqr 11.6 Aqr 11.9 Cet 12.1 onal offset fixe	Max H (dec 40N 22 24 25 26 27 28 29 29 29 ed at T+60 c	El g) 40S 48 45 42 40 37 35 33 31 days)	
			6P/d'Arre	est			
Mag 9 10 11 12 13 14 15 2021	2021		2021	2021	2021	2022	Date
July 1	Aug.1	Sept.1	Oct.1 1	Nov.1 I	Dec.1	Jan.1	Feb.1
Recent Magnitude Recent Magnitu Comet Des YYY None	e Measurements C de Measurement Y MM DD.DD (UT)	Contributed to the s in ICQ form Mag SC APER T	e ALPO Comets nat: FL POW CO Dia	<u>Section</u> MA TAIL DC LENG	ICQ CC PA	DDE Observer	Name

6P/d'Arrest is also past perihelion [T = 2021 September 17 @ 1.35 au]. Due to an asymmetrical lightcurve, the comet peaks in brightness a month or two after perihelion. December should see the comet fade from around magnitude 10.7 to 11.9 though observations in late October and early November suggest 6P may be 0.5-1.0 magnitudes brighter than the above prediction.

This month, 6P is an evening object moving through Pisces Austrinus (Dec 1-12), Aquarius (12-31) and Cetus (31). Though better placed for southern observers, it is observable from both hemispheres.

Discovered 1927 November 15 by the Arnold Schwassmann and Arno Arthur Wachmann at the Hamburg Observatory in Bergedorf, Germany

Centaur comet with orbital period of ~14.8 years

Orbit (from Minor Planet Center, MPEC 2021-W138)

29P/Schwassmann-Wachmann Epoch 2022 Jan. 21.0 TT = JDT 2459600.5 T 2019 Apr. 4.85113 TT Rudenko q 5.7713405 (2000.0) P Q n 0.06636470 Peri. 49.81465 +0.99174187 -0.04468182 a 6.0419613 Node 312.38189 -0.02059024 +0.86971579 e 0.0447902 Incl. 9.36627 +0.12658632 +0.49152618 P 14.9 From 11652 observations 2018 June 18-2021 Nov. 29, mean residual 0".6.

Ephemerides (produced with Seiichi Yoshida's Comets for Windows program)

29P/Schwassmann-Wachmann Ma												Мах	: El					
																	((leg)
I	Date		R	.A.	Dec	21.	r			d]	Elong	Сс	nst	Mag	4	ΟN	40S
2021	Dec	01	04	37	+31	59	5.9	35	4	4.963		169M	I	Per	10-13		82	18
2021	Dec	06	04	35	+31	52	5.9	37	4	4.964		170E	I	Per	10-13		82	18
2021	Dec	11	04	32	+31	44	5.9	38	4	4.973		167E	I	Per	10-13		82	18
2021	Dec	16	04	29	+31	34	5.9	40	4	4.990		163E	I	Per	10-13		82	18
2021	Dec	21	04	27	+31	25	5.9	41	ļ	5.014		158E	I	Per	10-13		81	19
2021	Dec	26	04	24	+31	14	5.9	43	!	5.045		153E	I	Per	10-13		81	19
2021	Dec	31	04	22	+31	03	5.9	44	ļ	5.084		148E	I	Per	10-13		81	19
2022	Jan	05	04	20	+30	52	5.9	46	ļ	5.129		143E	1	lau	10-13		81	19

Comet Magnitude Formula

None, due to frequent outbursts.

Recent Magnitude Measurements Contributed to the ALPO Comets Section

Recent Mag	Recent Magnitude Measurements in ICQ format:										
Comet Des	YYYY MM DD.DD	Mag SC APER FL POW	COMA	TAIL ICQ	CODE	Observer Name					
	(UT)	Т	Dia DC	LENG PA							
29	2021 12 02.77 S 1	1.3 TI 29.8L 4 108	3 1	ICQ X	X HAR11	Christian Harder					
29	2021 11 13.69 xS 1	1.0 AQ 40.0L 4 59	2.9 2/	ICQ X	X WYA	Christopher Wyatt					
29	2021 11 11.09 S 1	0.1 TK 20.3T10 77	5 3	ICQ X	X GON05	Juan Jose Gonzalez Suarez					
29	2021 11 09.92 S 1	0.8 TI 29.8L 4 92	4 2	ICQ X	X HAR11	Christian Harder					
29	2021 11 05.11 S 1	0.0 TK 20.3T10 77	4 3	ICQ X	X GON05	Juan Jose Gonzalez Suarez					
29	2021 11 04.50 S 1	0.7 TK 12.5B 30	3 2	ICQ x	x HER02	Carl Hergenrother					
29	2021 11 02.13 S 1	0.6 TI 29.8L 4 92	2.9 2	ICQ X	X HAR11	Christian Harder					
29	2021 11 01.85 S 1	1.4:TI 53.1L 139	1.8 3	ICQ X	X HAR11	Christian Harder					

29P/Schwassmann-Wachmann was discovered photographically on 1927 November 15 by German observing team Arnold Schwassmann and Arno Arthur Wachmann. 29P is one of the more enigmatic comets as it experiences outbursts multiple times per year that can reach 10-14th magnitude.

29P has been especially active of late with multiple outbursts observed since late September. As a result, the comet is about as bright as it ever gets with many visual observers reporting the comet to be between magnitude 10.0 and 11.4 with a coma diameter between \sim 2-5'.

The comet is at opposition on December 2 in Perseus and observable from both hemispheres. If you observe 29P, please consider contributing to two pro-am efforts to better understand this object: the British Astronomical Society's (BAA) Mission 29P monitoring program coordinated by Richard Miles. (<u>https://britastro.org/node/18562</u> & <u>https://britastro.org/node/25120</u>) and the University of Maryland's 29P Observation campaign (<u>https://wirtanen.astro.umd.edu/29P/29P_obs.shtml</u>).

57P/du Toit-Neujmin-Delporte

Discovered on 1941 July 18 by Daniel du Toit at the Harvard College Observatory's Boyden Station in South Africa, on 1941 July 25 by Grigory N. Neujmin at the Simeis Observatory in Russia, and on 1941 August 19 by Eugéne Joseph Delporte of the Royal Observatory in Uccle, Belgium Jupiter-family comet with orbital period of 6.4 years Orbit (from Minor Planet Center, MPEC 2021-W138) 57P/du Toit-Neujmin-Delporte Epoch 2022 Jan. 21.0 TT = JDT 2459600.5T 2021 Oct. 17.39693 TT Rudenko (2000.0)1.7200339 Peri. 115.25496 Node 188.76828 Incl. 2.85132 Ρ 0 q +0.82889115 n 0.15397450 +0.55935853 3.4474940 0.5010770 а -0.77729950 +0.52132122 -0.28796444 +0.20288827е Ρ 6.40 From 1126 observations 2015 Feb. 18-2021 Nov. 29, mean residual 0".7. Ephemerides (produced with Seiichi Yoshida's Comets for Windows program) 57P/du Toit-Neujmin-Delporte Max El (deq) d 40N 40S Date R.A. Decl. r Elong Const Mag 2021 Dec 01 20 12 -18 24 1.769 2.187 52E 50E 21 Cap 11-13 17 2021 Dec 06 20 26 -17 44 1.780 2.231 Cap 11-13 21 14 Cap 11-13 2021 Dec 11 20 40 -17 00 1.792 2.276 49E 21 11 2021 Dec 16 20 53 -16 12 1.806 2.321 47E Cap 11-13 21 8 2021 Dec 21 21 07 -15 21 1.820 2.367 45E Cap 11-13 20 6 2021 Dec 26 21 20 -14 28 1.835 2.414 43E Agr 11-13 20 4 2021 Dec 31 21 33 -13 31 1.851 2.461 41E Cap 11-13 19 2 2022 Jan 05 21 46 -12 32 1.868 2.509 40E Cap 11-13 18 0 Comet Magnitude Formula Currently in outburst Recent Magnitude Measurements Contributed to the ALPO Comets Section Recent Magnitude Measurements in ICQ format: Comet Des YYYY MM DD.DD Mag SC APER FL POW COMA TAIL ICO CODE Observer Name (UT) Т Dia DC LENG PA 2021 11 09.71 S 11.0:TK 32.0L 5 80 57 0.5 TCO XX PTL01 Uwe Pilz

Daniel du Toit was the first to discover 57P/du Toit-Neujmin-Delporte on 1941 July 18 from the Harvard College Observatory's Boyden Station in South Africa only a few days after a close approach to Earth of 0.30 au. Due to World War II, communications were slow and two other observers, Grigory N. Neujmin at Simeis Observatory in Russia and Eugéne Joseph Delporte of the Royal Observatory in Uccle, Belgium also found the comet over the next month or so. 57P is making its 9th observed return and was not expected to become much brighter than 16th magnitude. That was the case until October 17, its perihelion date, when it was observed 5 magnitudes brighter at 11th magnitude.

While not as outburst prone as 29P, 57P experienced a 6-magnitude outburst in 1996 which may have produced 19 or more secondary nuclei that were observed during its next return in 2002. Its abnormal brightness in 1941 also suggests an outburst in that year.

While observations to the ALPO and COBS have been few in November, they do suggest the comet is still as bright as magnitude 11.0 on November 9 (Uwe Pilz) and 11.7 on November 24 (Steffen Fritsche to COBS). December sees 57P as an evening object in Capricornus (Dec 1-25), Aquarius (25-28), and Capricornus (28-31). Unless another outburst occurs, 57P should fade over the coming weeks.

Discovered 2017 May 21 by the Pan-STARRS survey with the Pan-STARRS1 1.8-m on Haleakala Dynamically old long-period comet

Orbit (from MPEC 2021-U138)

C/2017 K2 (PAN	STARRS)				
Epoch 2022 Jan. 21	.0 TT = .	JDT 2459600.5			
T 2022 Dec. 19.691	96 TT			Rudenko	
q 1.7971155	(2	2000.0)	P	Q	
z -0.0003894	Peri. 2	236.19330	+0.0182544	+0.0492546	52
+/-0.000008	Node	88.23673	-0.1810170	+0.9824431	4
e 1.0006999	Incl.	87.55886	-0.983310	-0.1799429)5
From 7072 observat	ions 201	3 May 12-2021	Nov. 22, r	mean residual 0".	4.
1/a(orig) = -0.000	031 AU**	-1, 1/a(fut) =	+0.001162	L AU**-1.	

Ephemerides (produced with Seiichi Yoshida's Comets for Windows program)

C/2017 K2 (PANSTARRS) Max												
								(d	eg)			
Date	R.A.	Decl.	r	d	Elong	Const	Mag	40N	40S			
2021 Dec 01	17 33	+16 20	4.671	5.366	41E	Her	11.6	19	0			
2021 Dec 06	17 37	+15 44	4.627	5.339	39E	Her	11.6	16	0			
2021 Dec 11	17 41	+15 11	4.582	5.308	38E	Her	11.5	12	0			
2021 Dec 16	17 45	+14 40	4.538	5.272	38E	Her	11.5	9	0			
2021 Dec 21	17 49	+14 12	4.494	5.232	37M	Oph	11.4	8	0			
2021 Dec 26	17 53	+13 46	4.449	5.187	37M	Oph	11.4	11	0			
2021 Dec 31	17 57	+13 22	4.405	5.138	37M	Oph	11.3	14	0			
2022 Jan 05	18 01	+13 00	4.360	5.084	38M	Oph	11.2	17	0			

Comet Magnitude Formula (from ALPO and COBS data)





C/2017 K2 (PANSTARRS) was discovered on 2017 May 21 by the Pan-STARRS1 1.8-m telescope at Haleakala on the Hawaiian island of Maui. At discovery the comet was around 21st magnitude and located at 16.1 au from the Sun. Pre-discovery observations were found back to May of 2013 when the comet was 23.7 au from the Sun which is further than the distance of Uranus. Even though it was discovered over 2.5 years ago, perihelion is still over a year away on 2022 December 19 at 1.80 au.

C/2017 K2 is poorly placed for observation this month as it passes through solar conjunction, though it will be ~37 degrees north of the Sun at conjunction. As a result, it will be invisible to southern hemisphere observers but visible at low elevations from the northern hemisphere (moving through Hercules [Dec 1-20] and

Ophiuchus [20-31]. Northern observers with a clear and dark northern horizon should be able to watch K2 brighten from around magnitude 11.6 to 11.2 this month.

The comet will reappear for southern hemisphere observers in February 2022 when it should be magnitude 10.5. Northern observers will be able to follow the comet continuously till late September when it will travel too far south (around magnitude 7.0 at that time). C/2017 K2 should peak in January 2023 around magnitude 6.5 and at a far southern declination of -70 deg. Northern observers won't see the comet again till August 2023 when it will have faded to around magnitude 10.0.

Like 8P/Tuttle, 19P/Borrelly, and C/2021 A1 (Leonard), C/2017 K2 will also have an orbit plane crossing this month (on December 19). While first time Oort Cloud comets usually have yet to release enough dust to produce dust trails before perihelion, K2 has been active out to 23 au and possible further. Imagers should attempt to image any dust trail in mid-December since the existence and morphology of a trail may provide additional constraints on the start of K2's activity.

Somee du 8 au 3. oct S'firis - 2021. Lusouie - Michel. Luppuie 2017 C P. HERCULE. Jush 45 TU 460m. O Sma 204× Centre + marqué

Figure 13 – Michel Besson and Ludovic Prebet sketched C/2017 K2 (PANSTARRS) on 2021 October 8.

New Comet Numberings (Ref: WGSBN Bull. 1 #12)

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436P/2007 R4
               = 2021 U2 (Garradd)
435Р/2021 ТЗ
               = 2015 \text{ K6} (PANSTARRS)
434P/2012 TK8
               = 2021 S2 (Tenagra)
433P = (248370) Dual status
               = P/2016 U2 (PANSTARRS)
432P/2021 N4
431P/2015 Q1
               = P/2021 P5 (Scotti)
430P/2011 A2
               = P/2021 Q2 (Scotti)
429P/2008 QP20 = P/2021 M1 (LINEAR-Hill)
428P/2014 W12 = P/2021 Q1 (Gibbs)
427P/2017 S5
               = P/2021 L6 (ATLAS)
426P/2019 A7
               = P/2021 K4 (PANSTARRS)
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New Comet Discoveries

P/2021 V3 = P/2011 UE215 (PANSTARRS) - Discovered by the Pan-STARRS survey at 21st magnitude withthe Pan-STARRS1 1.8-m on 2021 November 1. After additional observations in 2021 were found, the MinorPlanet Center linked the new comet with observations in 2012 and a designated object from 2011 called 2011UE215. Perihelion will be on 2022 August 18 at 3.40 au. According to Syuichi Nakano, the comet passed 0.39au from Jupiter in 2007. Prior to the 2007 encounter, its perihelion was larger at 3.95 au. P/2021 V3 is unlikelyto get brighter at this return. [CBET 5069, MPEC 2021-V173]

P/2021 V2 (Fuls) – Discovered by David Carson Fuls at 19th magnitude with the Mount Lemmon 1.5-m on 2021 November 7. This comet is a short-period comet with an orbital period of 27.2 years. Perihelion is on 2023 January 21 at 3.50 au. Perihelion will be next year on 2022 April 30 at 3.01 au with a peak brightness of 17th magnitude. [CBET 5068, MPEC 2021-V169]

C/2021 V1 (Rankin) – Discovered by David Rankin at 20th magnitude with the Mount Lemmon 1.5-m on 2021 November 5. Perihelion will be next year on 2022 April 30 at 3.01 au. The comet is unlikely to get brighter than 19th magnitude.[CBET 5067, MPEC 2021-V167]

C/2021 U5 (Catalina) – The Catalina Sky Survey found yet another comet. This time it was Alex Gibbs who used the Catalina 0.68-m Schmidt to find this 18th magnitude comet on October 29. Perihelion will be next month on 2022 January 25 at 2.37 au. It should peak in brightness next February and March at ~17th magnitude. [CBET 5070, MPEC 2021-V199]

As always, the Comet Section is happy to receive all comet observations, whether textual descriptions, images, drawings, magnitude estimates, or spectra. Please send your observations via email to the Comets Section < comets @ alpo-astronomy .org >, Comets Section Coordinator Carl Hergenrother < carl.hergenrother @ alpo-astronomy .org > and/or Comets Section Acting Assistant Coordinator Michel Deconinck < michel.deconinck @ alpo-astronomy .org >.

Thank you to everyone who contributed to the ALPO Comets Section!

Clear skies! - Carl Hergenrother 26

Recent Magnitudes Contributed to the ALPO Comets Section

Comet Des	YYYY MM	DD.	DD	Mag	r SC	APER 1	FL POW	COM	A	TAIL			ICÇ	CODE	Observer Name	
C/2021 M1	U)	T)				Т		Dia	DC	LENG	PA					
2021 AI	2021 12	02	49 7.	65	114	7 2R	5a600	26 3	>	1 0 3	313	TCO	vv	HEB02	Carl Hergeprother	
2021A1	2021 12	02.	48 M	6.7	' TK	5.0B	10	8	5/	0.7 3	320	TCO	XX	HER02	Carl Hergenrother	
2021A1	2021 12	01.	46 M	7.0	TK	5.0B	10	6	5/	0.3 3	320	ICQ	хх	HER02	Carl Hergenrother	
2021A1	2021 12	01.	17 I	6.8	TK	12.6B	5 25	10	6	45.Om3	300	ICQ	XX	DECaa	Michel Deconinck	
2021A1	2021 11	30.	47 M	7.1	TK	5.0B	10	5	5	0.3 3	320	ICQ	XX	HER02	Carl Hergenrother	
2021A1	2021 11	28.	50 M	7.2	TK	5.0B	10	7	5	0.2 3	315	ICQ	ХХ	HER02	Carl Hergenrother	
2021A1	2021 11	28.	18 E	7.9	TK	25.0C	10 62	6	4	25.0m3	300	ICQ	XX	DECaa	Michel Deconinck	
2021A1	2021 11	26.	46 C	7.9	GG	5.0R	4a 60	10.6	F /	- 0 -	. 1 0	ICQ .	XX	OLAaa	Michael Olason	
2021A1 2021A1	2021 11	10	15 S 24 g	8.4	TK	32.0L	5 8U 25	4	2/ 2	0 15 3	370	TCQ .	XX VV	CON05	UWE FILZ	Suaroz
2021A1 2021A1	2021 11	17 17	24 D 51 Q	8 2	י דת יידיג	10.0B	2.J 1.0	0 7	2 2	0.13 3	520	TCO	AA VV	GONUJ HERO2	Carl Hergeprother	Suarez
2021A1	2021 11	17.	50 V	8.5	U4	10.6R	5a600	13.4	-	20 m.3	320	TCO	XX	HER02	Carl Hergenrother	
2021A1	2021 11	17.	47 k	9.0	U4	10.6R	5a600	8.2		26 m3	325	ICO	xx	HER02	Carl Hergenrother	
2021A1	2021 11	16.	20 I	8.4	ΤK	25.0C	10 62	2.5	7	5.0m2	280	ICQ	XX	DECaa	Michel Deconinck	
2021A1	2021 11	15.	50 k	9.3	U4	10.6R	5a360	5.8		15 m3	325	ICQ	XX	HER02	Carl Hergenrother	
2021A1	2021 11	15.	48 V	9.0	U4	10.6R	5a600	12.8		12 m3	320	ICQ	ΧХ	HER02	Carl Hergenrother	
2021A1	2021 11	15.	52 S	8.2	ΤK	5.0B	10	5	4			ICQ	XX	HER02	Carl Hergenrother	
2021A1	2021 11	15.	51 M	8.4	TK	12.5B	30	5	6	6 m3	310	ICQ	ХХ	HER02	Carl Hergenrother	
2021A1	2021 11	13.	50 S	8.5	TK	5.0B	10	6	4	E		ICQ :	XX	HER02	Carl Hergenrother	
2021A1	2021 11	13.	50 M	. 8.9	I'K	12.5B	30	5	6	5 m.:	320	TCQ	XX 	HERU2	Carl Hergenrother	
2021A1 2021A1	2021 11	11	51 S 22 g	9.4	TK.	12.5B	30	с о	2	0 1 3	320	TCO	XX VV	CON05	Tuan Jose Consales	Suaroz
2021A1 2021A1	2021 11	11	22 3 21 9	9.2	11. TR	20 3TT	10 77	6	3/ 3/	0.4 .	320	TCO	AA VV	GON05	Juan Jose Gonzalez	Suarez
2021A1	2021 11	10.	17 S	9.4	TT	20.31. 29.8L	4 92	4	4	6.5m ⁻	310	TCO	XX	HAR11	Christian Harder	Duurez
2021A1	2021 11	07.	49 S	9.9	TK	12.5B	30	3	5	0.01.0		ICO	XX	HER02	Carl Hergenrother	
2021A1	2021 11	05.	23 S	9.7	ΤK	20.3T	10 77	5	3/	0.3 3	330	ICQ	XX	GON05	Juan Jose Gonzalez	Suarez
2021A1	2021 11	04.	49 S	10.0	TK	12.5B	30	4	4			ICQ	XX	HER02	Carl Hergenrother	
2021A1	2021 11	02.	15 S	10.3	TI	29.8L	4 79	3	4	6.Om3	340	ICQ :	XX	HAR11	Christian Harder	
C/2019 L3	(ATLAS)															
2019L3	2021 12	02.	78 S	9.7	ΤI	29.8L	4 92	1.9	4/			ICQ :	XX	HAR11	Christian Harder	
2019L3	2021 11	13.	'/2 xM	10.2	AQ	40.0L	4 59	4	6	5.8m2	290	ICQ	XX	WYA	Christopher Wyatt	~
2019L3	2021 11		13 5	9.6) TK	20.3T.	10 //	4	5	4 5		TCQ	XX	GON05	Juan Jose Gonzalez	Suarez
201913	2021 11	09.	93 5 35 0	10.2	. TI 1. mizi	29.8L 12 5D	4 92	2.4	2	4.5003	505	TCQ .	AA VV	HARII UFD02	Carl Wargaprother	
2019L3 2019L3	2021 11	05.	55 5 13 5	9.7	TK	20 3T	10 77	5	4			TCO	XX XX	GON05	Juan Jose Conzalez	Suarez
2019L3	2021 11	02.	12 S	10.2	TT	29.8T	4 79	1.8	4	6.0m2	275	TCO	XX	HAR11	Christian Harder	Duurez
C/2017 K2	(PANSTAR	RS)										2				
2017K2	2021 11	01.	79 S	11.6	TI 🗧	53.1L	155	1.3	3			ICQ	XX	HAR11	Christian Harder	
241P/LINEA	R															
241	2021 11	15.	42 C	17.2	GG	27.9T	6A200	0.2		> 0.3m2	270	ICQ :	XX	OLAaa	Michael Olason	
173P/Muell	er	1 -	45 0	10.0		07 07	67.600	0 0		1 0 0						
1220/11-1-	ZUZI II Demon	15.	45 C	19.8	GG	27.91	6A68U	0.2		1.0m2	279	ICQ .	XX	OLAaa	Michael Olason	
132P/HeIIN	2021 11	11	02 G	11 0	אידי	20 300	10 133	з	Л			TCO	vv	CON05	Juan Jose Conzalez	Suaroz
132	2021 11	01	81 S	14 5	, IN TT	53 1T.	215	05	4			TCO	XX	HAR11	Christian Harder	Suarez
116P/Wild	2021 11	01.	OT D	11.0		55.11	210	0.0	-			TOŽ	2121	111111111	CHILISCIAN NALACI	
116	2021 11	15.	43 C	16.6	GG	27.9T	6a900	0.2	:	> 1.5m2	285	ICQ	XX	OLAaa	Michael Olason	
110P/Hartl	ey															
110	2021 11	15.	39 C	15.0	GG	27.9T	6a900	0.3		> 1.3m2	287	ICQ	XX	OLAaa	Michael Olason	
106P/Schus	ter															
106	2021 11	15.	34 C	17.1	GG	27.9T	6a900	0.2		> 2 m2	281	ICQ :	XX	OLAaa	Michael Olason	
70P/Kojima		1 -		1 7 0		07 07	c 000	0 0								
/U	2021 11	15.	49 C	1/.2	GG	27.91	6a900	0.2		> 0.5m2	299	ICQ .	XX	OLAaa	Michael Olason	
67 67	2021 11	asıme 15	enko 35 C	9.2	CC	27 0.07	6-910					TCO	vv		Michael Olacon	
67	2021 11	13	55 € 67 ⊽M	9.2	UGU. TR	40 OT.	0 a 0 4 0 4 5 9	33	6	12 0m2	275	TCO	AA VV	WYA WYA	Christopher Wyatt	
67	2021 11	11.	11 S	9.2	 	10.0B	25	4	5/	12.01112		TCO	XX	GON05	Juan Jose Gonzalez	Suarez
67	2021 11	11.	10 S	9.8	TK	20.3T	10 77	4	5	0.25 2	280	ICO	XX	GON05	Juan Jose Gonzalez	Suarez
67	2021 11	09.	94 S	9.5	TI	29.8L	4 92	2	4	10.0m2	280	ICQ	XX	HAR11	Christian Harder	
67	2021 11	07.	21 B	9.6	TK	25.0C	10 195	1.8	7	2.2m2	260	ICQ	XX	DECaa	Michel Deconinck	
67	2021 11	05.	35 S	9.0	TK	12.5B	30	4	6/			ICQ	XX	HER02	Carl Hergenrother	
67	2021 11	05.	21 S	9.9	TK	20.3T	10 77	4	5	0.15 2	280	ICQ	XX	GON05	Juan Jose Gonzalez	Suarez
67	2021 11	02.	11 S	9.2	TI	29.8L	4 79	2.3	4/	11.Om2	280	ICQ	XX	HAR11	Christian Harder	
57P/du Toi	t-Neujmi	n-De	⊥port	e		20 0-	F 00	0 5				T.C.C.		DTT 01		
5/ 52D/U222-	2021 11 aton %ha	. 09.' 	/⊥ S	11.0	:'1'K	32.0L	5 80	0.5				TCŐ :	ΧХ	RIT01	UWE PILZ	
52 JZF/Harrin	90011-ADE	:⊥⊥ 15	53 C	17 0	CC.	27 0.07	62720	0 2		> 0 0m2	рал	TCO	vv	OT.D > >	Michael Olason	
JZ	2021 11	т).	JJ (± / • 0	90	21.91	Ja / 2 U	0.2		- 0.9112	- 74	TON .	ΛΛ	JUAAd	MICHAEL ULASUII	

29P/Schwass	smann-	-Wad	chmann													
29	2021	12	02.77	S	11.3	ΤI	29.8L 4	108	3	1			ICQ	XX	HAR11	Christian Harder
29	2021	11	13.69	xS	11.0	AQ	40.0L 4	59	2.9	2/			ICQ	XX	WYA	Christopher Wyatt
29	2021	11	11.09	S	10.1	ΤK	20.3T10	77	5	3			ICQ	XX	GON05	Juan Jose Gonzalez Suarez
29	2021	11	09.92	S	10.8	ΤI	29.8L 4	92	4	2			ICQ	XX	HAR11	Christian Harder
29	2021	11	05.11	S	10.0	ΤK	20.3T10	77	4	3			ICQ	XX	GON05	Juan Jose Gonzalez Suarez
29	2021	11	04.50	S	10.7	ΤK	12.5B	30	3	2			ICQ	XX	HER02	Carl Hergenrother
29	2021	11	02.13	S	10.6	ΤI	29.8L 4	92	2.9	2			ICQ	XX	HAR11	Christian Harder
29	2021	11	01.85	S	11.4	:TI	53.1L	139	1.8	3			ICQ	XX	HAR11	Christian Harder15P/Finlay
15	2021	11	15.40	С	16.9	GG	27.9T 6a	1900	0.4		>	0.5m285	ICQ	XX	OLAaa	Michael Olason
9P/Tempel																
9	2021	11	15.54	С	14.8	GG	27.9T 6a	1300	0.4				ICQ	XX	OLAaa	Michael Olason
8P/Tuttle																
8	2021	11	13.71	xS	10.4	ΤK	40.0L 4	59	3.5	3			ICQ	XX	WYA	Christopher Wyatt
4P/Faye																
4	2021	11	13.72	xS	11.6	AQ	40.0L 4	108	1.9	3			ICQ	XX	WYA	Christopher Wyatt
4	2021	11	11.17	S	11.0	ΤK	20.3T10	100	3	2/			ICQ	XX	GON05	Juan Jose Gonzalez Suarez
4	2021	11	10.19	S	10.3	ΤI	29.8L 4	92	2.8	2			ICQ	XX	HAR11	Christian Harder
4	2021	11	05.36	S	10.6	ΤK	12.5B	30	2	4			ICQ	хx	HER02	Carl Hergenrother
4	2021	11	05.09	S	11.2	ΤK	20.3T10	100	3	2/			ICQ	XX	GON05	Juan Jose Gonzalez Suarez