## **Meteor Activity Outlook for October 5-11, 2024**

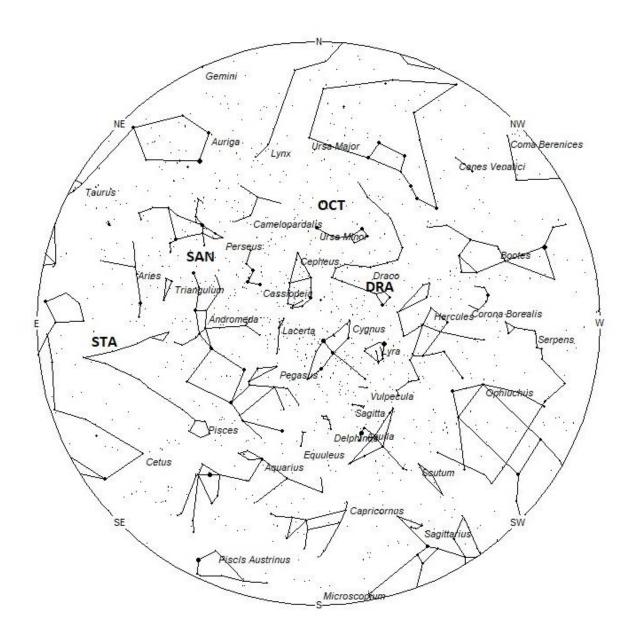


Ryan Conner captured this brilliant fireball using his <u>AllSky Camera System</u> on May 1, 2024, at 3:46 EDT (7:27 UT) from North Royalton, Ohio USA. The secondary streak is an artifact due to the extreme brightness of this fireball. ©Ryan Conner

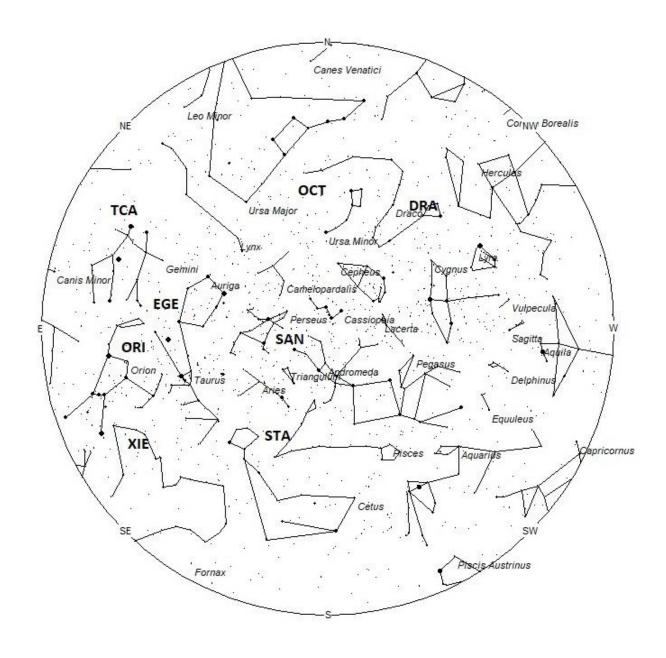
During this period, the moon reaches its first quarter phase on Thursday October 10th. At that time the moon will lie 90 degrees east of the sun and will set between 23:00 and midnight local daylight-saving time (LDST) on that night. This weekend the waxing crescent moon will set during the early evening hours and will set long before the more active morning hours arrive. The estimated total hourly rates for evening observers this weekend should be near 3 as seen from mid-northern latitudes (45N) and 2 as seen from tropical southern locations (25S) For morning observers, the estimated total hourly rates should be near 13 as seen from mid-northern latitudes (45N) and 10 as seen from tropical southern locations (25S). The actual rates seen will also depend on factors such as personal light and motion perception, local weather conditions, alertness, and experience in watching meteor activity. Note that the hourly rates listed below are estimates as viewed from dark sky sites away from urban light sources. Observers viewing from urban areas will see less activity as only the brighter meteors will be visible from such locations.

The radiant (the area of the sky where meteors appear to shoot from) positions and rates listed below are exact for Saturday night/Sunday morning October 5/6. These positions do not change greatly day to day so the listed coordinates may be used during this entire period. Most star atlases (available at science stores and planetariums) will provide maps with grid lines of the celestial coordinates so that you may find out exactly where these positions are located in the sky. I have also included charts of the sky that display the radiant positions for evening, midnight, and morning. The center of each chart is the sky directly overhead at the appropriate hour. These charts are oriented for facing south but can be used for any direction by rotating the charts to the desired direction. A planisphere or computer planetarium program is also useful in showing the sky at any time of night on any date of the year. Activity from each radiant is best seen when it is positioned highest in the sky, either due north or south along the meridian, depending on your latitude. Radiants that rise after midnight will not reach their highest point in the sky until daylight. For these radiants, it is best to view them during the last few hours before dawn. It must be remembered that meteor

activity is rarely seen at its radiant position. Rather they shoot outwards from the radiant, so it is best to center your field of view so that the radiant lies toward the edge and not the center. Viewing there will allow you to easily trace the path of each meteor back to the radiant (if it is a shower member) or in another direction if it is sporadic. Meteor activity is not seen from radiants that are located far below the horizon. The positions below are listed in a west to east manner in order of right ascension (celestial longitude). The positions listed first are located further west therefore are accessible earlier in the night while those listed further down the list rise later in the night.



Radiant Positions at 21:00 Local Daylight-Saving Time



Radiant Positions at 01:00 Local Daylight-Saving Time



Radiant Positions at 05:00 Local Daylight-Saving Time

## These sources of meteoric activity are expected to be active this week

The **Draconids** (**DRA**) are expected to be active on October 8/9 this year. The radiant lies at 17:32 (263) +56. This position lies in southeastern Draco, very close to the position occupied by the faint star known as Kuma (nu Draconis). These meteors are best seen just as soon as it becomes dark on the evening of October 8. With an entry velocity of 21 km/sec., the average Draconid meteor would be of slow velocity. These meteors are poorly seen from the southern hemisphere due to the low radiant altitude seen from below the equator.

We are now encountering inbound debris from comet 2P/Encke, which has its source superimposed upon the anthelion radiant. Since it has been shown that meteors from 2P/Encke are more numerous than the Anthelions, we will recognize this activity as the Southern Taurids (STA) from now until late November, when we no longer encounter remnants from comet 2P/Encke. Like the anthelion radiant, the source area is large and diffuse so observers can be liberal with the shower association of these meteors. Recent investigations of the Southern Taurids have revealed two clearly distinct components. The first component, also known as the October Arietids, represents the early and annual activity of Southern Taurids. It displays very little variation year to year. The latter component represents the main source of activity and is periodic. The early STA's are active from September 28 through November 2 and peaks on October 14<sup>th</sup>. The main component of the STA's are active from October 12 through November 27 and peaks on November 5th. The center of the large STA radiant is currently located at 01:59 (030) +08. This position lies in southeastern Pisces, 3 degrees southwest of the 4<sup>th</sup> magnitude star known as xi<sup>1</sup> Ceti. This radiant is best placed near 0100 LDST, when it lies on the meridian and is located highest in the northern sky. Rates at this time should be near 3 per hour no matter your location. With an entry velocity of 30 km/sec., the average STA meteor would be of medium-slow velocity.

The **62 Andromedids** (**SAN**) are another recent discovery by Dr. Peter Jenniskens using his video network called CAMS. These meteors are active on only eight nights centered on October 9<sup>th</sup>. On the 9<sup>th</sup>, the radiant is located at 02:32 (038) +47. This position lies in northeastern Andromeda, one degree northeast of the faint star known as 62 Andromedae. To best see these meteors, face toward the north near 02:00 LST. Rates at this time should be less than 1 per hour no matter your location. With an entry velocity of 17 km/sec., the average SAN meteor would be of slow velocity.

The **xi Eridanids** (**XIE**) are a recent discovery by Dr. Peter Jenniskens using his video network called CAMS. These meteors are active from October 5-16, peaking on the 11th. On that night, the radiant is located at 04:36 (069) -06. This position lies in northeastern Eridanus, 3 degrees west of the 4<sup>th</sup> magnitude star known as omega Eridani. To best see these meteors, face toward the north near 04:00 LST. Rates at this time should be less than 1 per hour no matter your location. With an entry velocity of 54 km/sec., the average XIE meteor would be of medium-fast velocity.

The **Orionids** (**ORI**) are active from October 2 through November 12, with maximum activity occurring on October 22nd. The radiant is currently located at 05:33 (083) +14, which places it in northern Orion, 4 degrees north of the 3rd magnitude star known as Meissa (lambda Orionis). This area of the sky is best placed for observing during the last dark hour prior to dawn, when it lies highest in the northern sky. Current rates are expected to be less than 1 per hour, no matter your location. With an entry velocity of 67 km/sec., the average ORI meteor would be of swift velocity.

The **epsilon Geminids** (**EGE**) are active from September 27<sup>th</sup> through November 7<sup>th</sup> with maximum activity occurring on October 16<sup>th</sup>. The radiant is currently located at 05:58 (090) +30. This area of the sky lies in southeastern Auriga, 3 degrees west of the 4th magnitude star known as kappa Aurigae. To best see these meteors, face toward the east during the last dark hour prior to dawn. Rates at this time should be less than 1 per hour no matter your location. With an entry velocity of 69 km/sec., the average EGE meteor would be of swift velocity.

The **tau Cancrids** (**TCA**) are a weak shower with a long activity period of seven weeks. They are active from September 27 through November 12 with maximum activity occurring on October 22nd. The radiant currently lies at 08:14 (124) +30, which places it in northwestern Cancer, 5 degrees northeast of the 1<sup>st</sup> magnitude star known as Pollux (beta Geminiorum). To best see these meteors face eastward during the last few hours of the morning prior to dawn. Expected hourly rates are less than 1 per hour no matter your location. With an entry velocity of 67 km/sec., the average TCA meteor would be of swift velocity.

The **Daytime Sextantids (DSX)** are active from September 22-October 13, with maximum activity occurring on October 1st. The current position of the radiant is 10:35 (159) -04. This position lies in central Sextans, 1 degree southeast of the faint star known as delta Sextantis. Rates at this time should be less than 1 per hour no matter your location. With an entry velocity of 32 km/sec., the average DSX meteor would be of medium-slow velocity. No matter your location, these meteors are difficult to observe as the radiant lies roughly 30 degrees from the sun. Therefore, these meteors may only be seen during the last hour prior to dawn, shooting upward from the eastern horizon.

The **October Camelopardalids (OCT)** are expected to peak near 16:00 UT on Saturday October 5<sup>th</sup>. These meteors are only visible for 32 hours, so your location is important to see these meteors. The best location to see the 2024 peak would be Alaska as these meteors are best seen during the last few hours prior to dawn. The radiant is located at 10:51 (163) +78. This position lies in a remote area of extreme northwestern Draco. See the charts for a better idea of this location. Note that this radiant is not listed on charts for the southern hemisphere as these meteors are not visible from south of the equator. Hourly rates are dependent on your location, but most of us are expected to see less than 1 per hour. With an entry velocity of 45 km/sec., the average OCT meteor would be of medium velocity.

**Sporadic** meteors are those meteors that cannot be associated with any known meteor shower. All meteor showers are evolving and disperse over time to the point where they are no longer recognizable. Away from the peaks of the major annual showers, these sporadic meteors make up the bulk of the activity seen each night. As seen from the mid-northern hemisphere (45N) one would expect to see during this period approximately 10 sporadic meteors per hour during the last hour before dawn as seen from rural observing sites. Evening rates would be near 3 per hour. As seen from the tropical southern latitudes (25S), morning rates would be near 7 per hour as seen from rural observing sites and 2 per hour during the evening hours. Locations between these two extremes would see activity between these listed figures.

The list below offers the information in tabular form of the showers that I feel are within reach of the visual observer to discern. Hourly rates are often less than one, so these sources are rarely listed as visual targets in most meteor shower lists. If you are like me and wish to associate as many

meteors as possible with known sources, then you will appreciate these listings. Before listing meteors from these obscure sources, you should attempt to prove these meteors belong to them and are not chance alignments of sporadic meteors. You can note parameters such as duration, length, radiant distance and the elevation of each meteor to help compute the probability of shower association. It should be remembered that slow meteors can be seen from fast showers, but fast meteors cannot be produced from slow showers. Slower showers are those with velocities less than 35/km per second. Slow meteors can appear from fast showers when they appear close to the radiant or low in the sky. The table located on page 22 of the IMO's 2024 Meteor Shower Calendar is a big help in aiding in the identification of meteors. If you record the length and duration of each meteor, you can use this chart to check the probability of the meteor belonging to a shower of known velocity. If the angular velocity is similar to the figure in the table, then your meteor probably belongs to that shower. Rates and positions are exact for Saturday night/Sunday morning.

SHOWER	DATE OF MAXIMUM ACTIVITY	CELESTIAL POSITION	ENTRY VELOCITY	CULMINATION	HOURLY RATE	CLASS
		RA (RA in Deg.) DEC	Km/Sec	Local Daylight- Saving Time	North- South	
Draconids (DRA)	Oct 09	17:32 (263) +56	21	18:00	1 - 1	III
Southern Taurids (STA)	Oct 14	01:59 (030) +08	30	02:00	3 - 3	II
62 Andromedids (SAN)	Oct 09	02:32 (038) +47	17	03:00	<1 -<1	IV
xi Eridanids (XIE)	Oct 11	04:36 (069)	54	05:00	<1 -<1	IV
Orionids (ORI)	Oct 22	05:33 (083) +14	67	06:00	<1 -<1	I
epsilon Geminids (SPE)	Oct 16	05:58 (090) +30	69	06:00	<1 -<1	II
tau Cancrids (TCA)	Oct 22	07:52 (118) +29	67	08:00	<1 -<1	IV
Daytime Sextantids (DSX)	Oct 01	10:35 (159) -04	32	11:00	<1 -<1	IV
October Camelopardalids (OCT)	Oct 05	10:51 (163) +78	45	11:00	<1 -<1	IV

You can keep track of the activity of these meteor showers as well as those beyond the limits of visual observing by visiting the <u>NASA Meteor Shower Portal</u>. You can move the sky globe to see different areas of the sky. Colored dots indicate shower meteors while white dots indicate

sporadic (random) activity. The large orange disk indicates the position of the sun so little activity will be seen in that area of the sky.

Class Explanation: A scale to group meteor showers by their intensity:

- Class I: the strongest annual showers with Zenith Hourly Rates normally ten or better.
- Class II: reliable minor showers with ZHR's normally two to ten.
- **Class III**: showers that do not provide annual activity. These showers are rarely active yet have the potential to produce a major display on occasion.
- Class IV: weak minor showers with ZHR's rarely exceeding two. The study of these showers is best left to experienced observers who use plotting and angular velocity estimates to determine shower association. These weak showers are also good targets for video and photographic work. Observers with less experience are urged to limit their shower associations to showers with a rating of I to III.